

**Advances in Plasma Turbulence**  
**after the 1972 publication**  
**of Davidson's book**  
**Methods in Nonlinear Plasma Theory**<sup>a</sup>

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# Introduction

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Ron's 1972 book

## Methods in Nonlinear Plasma Theory

was a milestone in plasma-physics publishing.

- It remains the cleanest exposition of plasma weak-turbulence theory (and many other things).
- It should be required reading for just about everybody (certainly present and future theorists).

## What has happened over the subsequent 35 years?

- Renormalized strong plasma turbulence theory (with many practical subtopics)
- Stochasticity and nonlinear dynamics as applied to plasmas
- Nonlinear gyrokinetic formalism
- Advanced computation

# Plasma turbulence theory dates from the 1960's.

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1959 – Kraichnan's direct-interaction approximation (DIA)

1960

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62 – Quasilinear theory (QLT)

63 – Multiple-time-scale methods (Frieman; Sandri);  
[Lorenz: Deterministic nonperiodic flow]

64 – Weak-turbulence theory (WTT)

65 – Kadomtsev's monograph

66 – Resonance-broadening theory (RBT) (Dupree; Weinstock)

67 – Orszag & Kraichnan: critique of RBT; plasma DIA

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69 – Chirikov's thesis (on stochasticity)

# The 1970's: Ron's Book; MSR Formalism; Stochasticity

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1970

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72 – **Davidson:** *Methods in Nonlinear Plasma Theory*

73 – Martin–Siggia–Rose formalism (classical renormalization)

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75 – Stochastic acceleration & heating (Smith & Kaufman; Karney)

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77

78 – Plasma DIA (DuBois; Krommes);

Treve's review: "Theory of chaotic motion with application to controlled fusion research"

79 – Littlejohn: "A guiding center Hamiltonian: A new approach"

# The Methodology of Weak Turbulence Theory

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For weak turbulence, the method of multiple time scales is very well suited (see Davidson):

$$\dot{\psi} = \underbrace{L}_{(i\Omega - \gamma)} \psi + \frac{1}{2} M \psi \psi. \quad (1)$$

Cumulants:  $C_1(t) \doteq \langle \psi \rangle(t)$ ,  $C_2(t, t') \doteq \langle \delta\psi(t) \delta\psi(t') \rangle$ , ... :

$$\partial_t C_1 = LC_1 + \frac{1}{2} M(C_1^2 + C_2), \quad (2a)$$

$$\partial_t C_2 = LC_2 + MC_1 C_2 + \frac{1}{2} MC_3, \quad (2b)$$

$$\partial_t C_3 = LC_3 + MC_1 C_3 + \frac{1}{2} M(2C_2^2 + C_4), \quad (2c)$$

⋮

Crudely, let  $\gamma/\Omega = O(\epsilon)$ . Then a consistent ordering is

$$C_1 \sim 1, \quad C_2 \sim \epsilon, \quad C_3 \sim \epsilon^2, \quad C_4 \sim \epsilon^3, \quad \dots \quad (3)$$

Expand in multiple time scales:

$$\begin{aligned}
 & \left( \frac{\partial}{\partial t_0} + \epsilon \frac{\partial}{\partial t_1} + \epsilon^2 \frac{\partial}{\partial t_2} + \dots \right) \left( \epsilon C_2^{(1)} + \epsilon^2 C_2^{(2)} + \dots \right) \\
 & \quad - 2\epsilon\gamma \left( \epsilon C_2^{(1)} + \epsilon^2 C_2^{(2)} \right) \\
 & = M \left( \epsilon^2 C_3^{(2)} + \epsilon^3 C_3^{(3)} + \dots \right), \tag{4a}
 \end{aligned}$$

$$\begin{aligned}
 & \left( \frac{\partial}{\partial t_0} + \epsilon \frac{\partial}{\partial t_1} + \dots \right) \left( \epsilon^2 C_3^{(2)} + \epsilon^3 C_3^{(3)} \right) \\
 & \quad + 3(i\Omega - \epsilon\gamma) \left( \epsilon^2 C_3^{(2)} + \dots \right) \\
 & = M \left[ 2 \left( \epsilon C_2^{(1)} + \epsilon^2 C_2^{(2)} \right)^2 + \epsilon^3 C_4^{(1)} + \dots \right]. \tag{4b}
 \end{aligned}$$

$O(\epsilon)$ :

$$\frac{\partial C_2^{(1)}(t_0, t_1, \dots)}{\partial t_0} = 0. \tag{5}$$

$O(\epsilon^2)$ :

$$\frac{\partial C_2^{(1)}}{\partial t_1} + \frac{\partial C_2^{(2)}}{\partial t_0} = 2\gamma C_2^{(1)} + M C_3^{(2)}, \quad (6a)$$

$$\frac{\partial C_3^{(2)}}{\partial t_0} + 3i\Omega C_3^{(2)} = M [C_2^{(1)}]^2. \quad (6b)$$

Note:  $3\Omega$  really means  $\Omega_{\vec{k}} + \Omega_{\vec{p}} + \Omega_{\vec{q}} \equiv \Delta\Omega$  (with  $\vec{k} + \vec{p} + \vec{q} = \vec{0}$ ).

$$C_3^{(2)} = \left( \frac{e^{-i\Delta\Omega t_0} - 1}{-i\Delta\Omega} \right) M [C_2^{(1)}]^2 \rightarrow \pi\delta(\Delta\Omega), \quad (7)$$

$$\left( \frac{\partial}{\partial t_1} - 2\gamma \right) C_2^{(1)}(t_1) + \underbrace{\frac{\partial C_2^{(2)}(t_0, t_1)}{\partial t_0}}_{\text{set to 0 to avoid secularity}} = M C_3^{(2)}(t_1). \quad (8)$$

set to 0  
to avoid secularity

We finally arrive (schematically) at the wave kinetic equation of WTT:

$$\left( \frac{\partial}{\partial t_1} - 2\gamma \right) C_2 = M^2 \delta(\Delta\Omega) C_2^2. \quad (9)$$

# The multiple-time-scale expansion breaks down for strong turbulence.

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In reality, the resonance is broadened:

$$\pi\delta(\Delta\Omega) \Rightarrow \operatorname{Re} \left( \frac{1}{-i(\Delta\Omega + i\Delta\eta)} \right) = \frac{\Delta\eta}{\Delta\Omega^2 + \Delta\eta^2} \quad (10a)$$

$$\approx \frac{\Delta\eta}{\Delta\Omega^2} + O(\Delta\eta^3) + O(\Delta\eta^5) + \dots \quad (10b)$$

Terms through **ALL ORDERS** are required to correctly account for the turbulent broadening. Note that perturbation theory *does not work* near the resonance  $\Delta\Omega = 0$ . **Renormalization is required** (see later discussion).

The order-by-order construction inherent in the multiple-time-scale expansion is not well suited to situations in which renormalization is essential.

# Resonances and secularities underlie the origins of stochasticity, chaos, and turbulence.

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Recall Hamiltonian dynamics and the problem of resonant denominators:

$$H(\vec{J}, \vec{\theta}) = H_0(\vec{J}) + \delta H(\vec{J}, \vec{\theta}). \quad (11)$$

Equations of motion:

$$\frac{\partial H}{\partial \vec{J}} = \dot{\vec{\theta}} = \vec{\Omega}(\vec{J}) + \frac{\partial \delta H}{\partial \vec{J}}, \quad (12a)$$

$$\frac{\partial H}{\partial \vec{\theta}} = -\dot{\vec{J}} = \frac{\partial \delta H}{\partial \vec{\theta}}. \quad (12b)$$

Let  $\delta H = \sum_{\vec{m}} e^{i\vec{m} \cdot \vec{\theta}} H_{\vec{m}}$ . Then with  $\vec{\theta} \approx \vec{\theta}_0 + \vec{\Omega}t$ ,

$$\Delta \vec{J} = - \int dt \frac{\partial \delta H}{\partial \vec{\theta}} \approx -i \sum_{\vec{m}} \frac{\vec{m} H_{\vec{m}} e^{i\vec{m} \cdot \vec{\theta}_0} (e^{i\vec{m} \cdot \vec{\Omega}t} - 1)}{\vec{m} \cdot \vec{\Omega}}. \quad (13)$$

Trouble near  $\vec{m} \cdot \vec{\Omega}(\vec{J}) = 0$  (see Kolmogorov–Arnold–Moser).

# Analogous problems occur in forced, dissipative systems; cf. strange attractors.

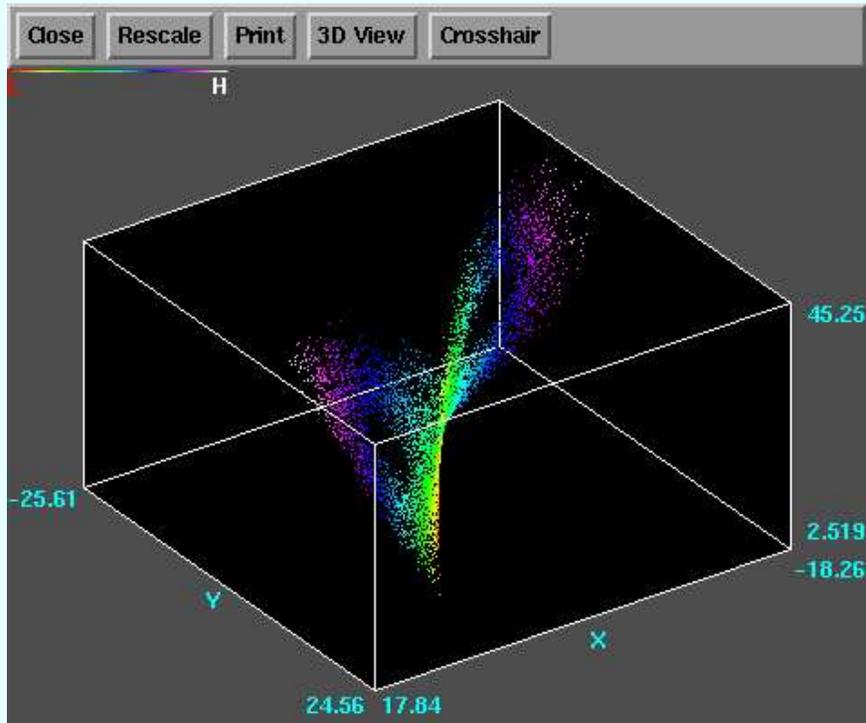


Fig. 1. The Lorenz strange attractor.

“butterfly effect”  $\Rightarrow$

- statistical description;
- cumulant hierarchy *through all orders*;
- renormalization.

**Stochastic / chaotic / turbulent systems exhibit exponential sensitivity to small changes in initial conditions.**

# A Very Brief Introduction to Renormalization

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A model of random passive advection illustrates the basic points:

$$\partial_t \psi(\vec{x}, t) + \vec{V}(\vec{x}, t) \cdot \vec{\nabla} \psi = 0, \quad (14a)$$

$$\partial_t \psi_{\vec{k}}(t) + i\tilde{\Omega}(t) \psi_{\vec{k}} = 0, \quad (14b)$$

where  $\tilde{\Omega}(t) \doteq \vec{k} \cdot \vec{V}(t)$ . Further, let  $\tilde{\Omega}(t) \rightarrow \tilde{\Omega}$ , a Gaussian random number with  $\langle \tilde{\Omega} \rangle = 0$ ,  $\langle \tilde{\Omega}^2 \rangle = \beta^2$ . Thus, study

$$\partial_t \psi = -i\tilde{\Omega} \psi. \quad (15)$$

The average response of the oscillator is characterized by the *mean Green's function (infinitesimal response function)*  $R(t; t')$ , which obeys in general

$$\partial_t R(t; t') + \int_{t'}^t d\bar{t} \underbrace{\Sigma(t; \bar{t})}_{\text{memory fcn. (turbulent coll op.)}} R(\bar{t}; t') = \delta(t - t'). \quad (16)$$

The goal is to calculate the memory function  $\Sigma$ .

# Linear dissipation is required to regularize an infinity.

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Less ambitiously, characterize the oscillator by a Markovian approx.,

$$\partial_\tau R(\tau) + \eta^{\text{nl}} R = \delta(\tau), \quad (17)$$

where  $\eta^{\text{nl}} \doteq \int_0^\infty d\tau \Sigma(\tau)$ , and try to calculate  $\eta^{\text{nl}}$ . In lowest-order perturbation theory (quasilinear level), it is easy to find that

$$\Sigma^{\text{QL}}(\tau) = \beta^2 R^{(0)}(\tau) \quad (\text{causal, but otherwise a constant}), \quad (18)$$

where  $R^{(0)}(\tau) = H(\tau)$  (unit step function). Thus

$$\eta^{\text{QL}} = \int_0^\infty d\tau \beta^2 \underbrace{R^{(0)}(\tau)}_1 = \begin{cases} 0 & (\beta = 0), \\ \infty & (\beta \neq 0). \end{cases} \quad (19)$$

To regularize the infinity, add some linear damping:

$$\partial_t \psi + \nu \psi = -i\tilde{\Omega} \psi. \quad (20)$$

Now  $R^{(0)}(\tau) \rightarrow H(\tau)e^{-\nu\tau}$ , and  $\eta^{\text{QL}} = \int_0^\infty d\tau \Sigma^{\text{QL}}(\rho) = \beta^2 / \nu$  (a continuous function of  $\beta$ ).

## Summing through all orders amounts to enforcing self-consistency — the essence of renormalization.

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With perturbation theory regularized, one can now proceed order by order, and eventually sum through all orders. The result is the replacement  $R^{(0)} \rightarrow R$ , i.e.,

$$\Sigma^{\text{nl}}(\tau) = \beta^2 R(\tau), \quad (21)$$

where  $R(\tau) = H(\tau) \exp[-(\nu + \eta^{\text{nl}})\tau]$ . Thus we obtain the self-consistent equation

$$\eta^{\text{nl}} = \frac{\beta^2}{\nu + \eta^{\text{nl}}}, \quad (22)$$

which can be solved for  $\eta^{\text{nl}}$ . In particular, one may now take the limit  $\nu \rightarrow 0$ , whereupon one finds  $\eta^{\text{nl}} = \beta$ . Note the appearance of *anomalous scaling*:

$$\eta^{\text{nl}} = \begin{cases} \beta^2/\nu & \text{(quasilinear theory),} \\ \beta^1 & \text{(renormalized).} \end{cases} \quad (23)$$

# Progress in Statistical Closure Theory

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1980

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82 – Terry & Horton: Drift-wave stochasticity;

Krommes: DIA for three-mode Terry–Horton equations

83 – Waltz: Numerical study of a Markovian closure;

[Lee: Gyrokinetic approach to particle simulation];

[Dubin<sup>a</sup> et al.: Hamiltonian approach to nonlinear gyrokinetic equations]

⋮ – More DIA and other closure calculations;

fluctuation noise in PIC simulations;

variational methods for bounding turbulent fluxes (Smith; Kim)

~1990 – **A major surprise and embarrassment: EDQNM closure doesn't work with waves (Bowman)**

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<sup>a</sup> Green  $\equiv$  Princeton graduate student.

# Progress, continued

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- 1993 – **Bowman** et al.: Realizable Markovian Closure (RMC)  
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- 95 – **Hu** et al.: Hasegawa–Wakatani equations and the RMC
- 96 – **[Holmes, Lumley, & Berkooz: *Turbulence, Coherent Structures, Dynamical Systems and Symmetry*]**
- 97
- 98 – **Diamond** et al.: Zonal flows and self-regulating DW turbulence  
99
- 2000 – **Krommes & Kim**: Interactions of disparate scales in drift-wave turbulence (unification of theories of ZFs and eddy viscosity)  
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:
- 2005 – **Nevins, Hammett, et al.**: Sampling noise in PIC simulations;  
**Kolesnikov**: “Bifurcation theory of the transition to collision-  
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less ion-temperature-gradient-driven plasma turbulence”

# An Example of State-of-the-Art Research on Plasma Turbulence: Blob Generation in a Turbulent Tokamak Edge

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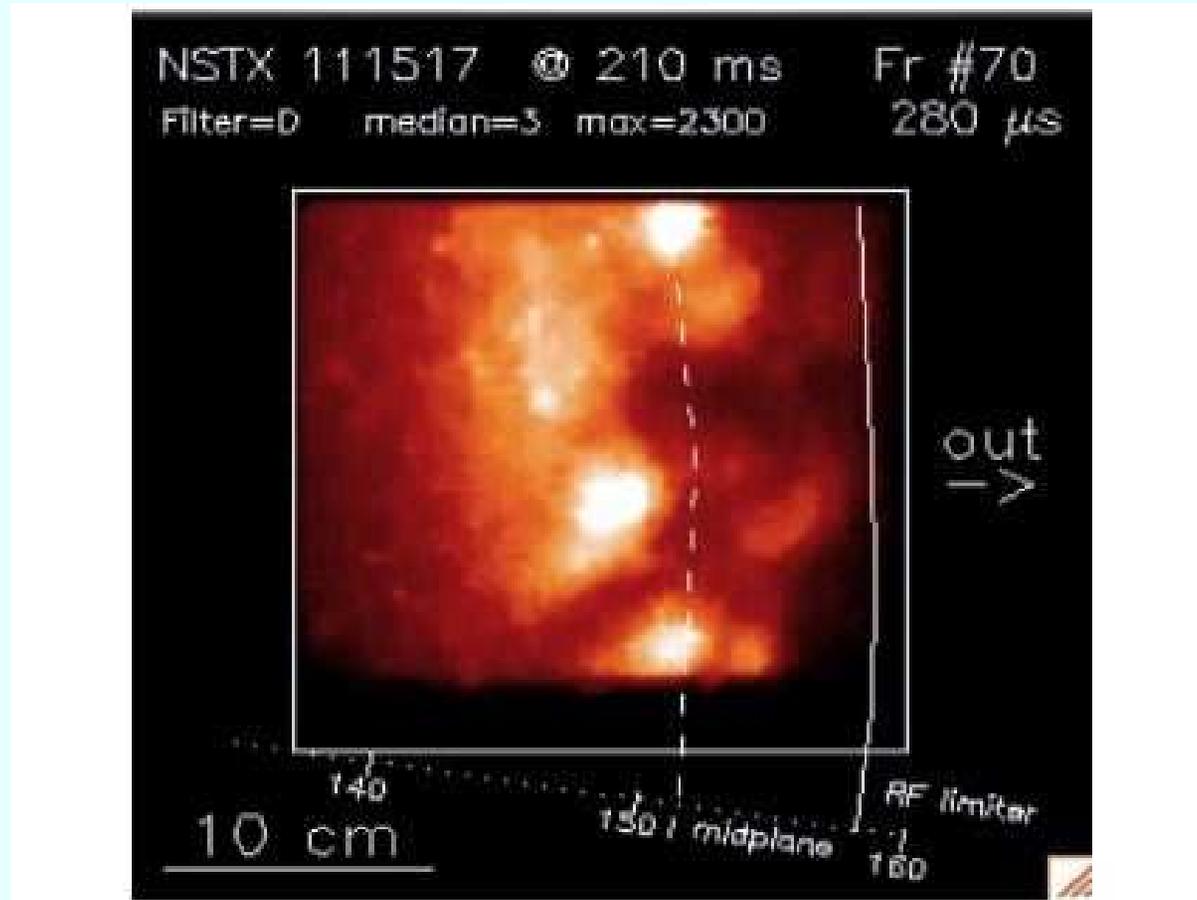


Fig. 2. Visualization of edge turbulence in NSTX using the gas-puff-imaging diagnostic. See <http://www.pppl.gov/~szweben/> for movies and more.

# The theory of blob generation requires knowledge and synthesis of many facets of linear and nonlinear physics.

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In the spirit of *Methods of Nonlinear Plasma Theory*, we would like to study blob generation analytically. This requires understanding of

- experimental phenomenology
- linear modes in a tokamak edge
- properties of turbulent saturated states
- nonlinear dynamics
- **intermittency**
- **formation of coherent structures**

**Analytical theories of the red topics  
are not well developed.**

# A research program to study blob generation theoretically is extremely challenging.

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- Derive paradigmatic nonlinear equation for NSTX edge.
- Numerically simulate to determine principal orthogonal eigenfunctions (good basis functions for the inhomogeneous turbulence).
- Galerkin project to deduce manageable system of coupled ODEs.
- Qualitative nonlinear-dynamics analysis (cycles in phase space describe birth–propagation–destruction of blobs).
- Estimate of blob generation rate
  - Understand dependence on important physical parameters.
  - Compare with experiment and simulations.
- **GRADUATE with PhD (Stoltzfus-Dueck)!**

**The ultimate goal is *dynamical control*.**

# To better train the next generation of plasma physicists, we need to rethink our curricula.

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Note: Append “for plasmas” to all topics below.

For everyone:

- more focus on nonlinear dynamics (appreciate the richness of possible behavior)
- fractal Brownian motion
- basics of fluid turbulence theory
- gyrokinetics (de-emphasize Braginskii)

For the hard-core theorists:

- Hamiltonian–Lie methods
- bifurcation theory
- serious dynamical-systems analysis
- renormalization techniques
- PDF methods
- intermittency
- coherent structures

These would be good chapters  
for Ron’s next book,  
*Methods in Nonlinear Plasma Theory, Vol. II.*

# Summary

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The organization of Ron's 1972 book applies just as well today, and highlights a challenging list of state-of-the-art problems in nonlinear plasma physics:

- Part I. Coherent Nonlinear Phenomena
  - $\vec{E} \times \vec{B}$  trapping
  - blob generation
- Part II. Turbulent Nonlinear Phenomena
  - zonal flows
  - anisotropic cascades in magnetic turbulence

**Ron's pioneering research and effective exposition have inspired multiple generations of scientists. He is a great role model.**

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